

# Influence of stratosphere-troposphere exchange on tropospheric ozone over the tropical Indian Ocean during the winter monsoon

M. Zachariasse,<sup>1</sup> P. F. J. van Velthoven,<sup>1</sup> H. G. J. Smit,<sup>2</sup>  
J. Lelieveld,<sup>3</sup> T. K. Mandal,<sup>4</sup> and H. Kelder<sup>1</sup>

**Abstract.** Ozone ( $O_3$ ) and relative humidity (RH) soundings, launched over the Indian Ocean during the 1998 winter monsoon (February–March), were analyzed. In the marine boundary layer (MBL),  $O_3$  mixing ratios were relatively low (10–20 ppbv) except close to the Indian subcontinent (40–50 ppbv) where profiles were strongly influenced by pollution. Sometimes, relatively low  $O_3$  levels were observed in the upper troposphere. These were associated with deep convection in regions where MBL  $O_3$  levels were also low. In the midtroposphere (300–500 hPa),  $O_3$  maxima (60–90 ppbv) were often found with low RH. A remarkable new finding of this study is that in more than a third of the profiles, laminae with very high  $O_3$  mixing ratios (up to 120 ppbv) were observed just below the tropical tropopause (between 100 and 200 hPa). Back trajectory analyses showed that these layers originated in the vicinity of the subtropical jet stream (STJ). We hypothesize that stratosphere-troposphere exchange (STE) near the subtropical jet by either shear-induced differential advection or clear-air turbulence (CAT) caused the midtropospheric maxima (STE followed by descent) and the upper tropospheric laminae. Another new finding is that stratospheric intrusions were not only found near the STJ but also deep within the tropics. Given the thickness of the midtroposphere intrusions (typically 3–5 km) and the very high  $O_3$  mixing ratios of the upper tropospheric laminae, it seems that STE plays an important role in the tropical tropospheric  $O_3$  budget, at least over the Indian Ocean during the winter monsoon.

## 1. Introduction

Ozone ( $O_3$ ) plays a key role in controlling the chemistry and climate of the tropical troposphere. Since it absorbs UV solar radiation, it is an important source for OH radicals. It is also an important greenhouse gas. Human activities increase the  $O_3$  concentration of the troposphere through emissions from fossil fuel combustion and biomass burning [Crutzen *et al.*, 1979, 1985; Logan *et al.*, 1981]. This will especially be the case in the future in the tropics where the economic activity is growing strongly with possible strong impact on chemistry and climate [Intergovernmental Panel on Climate (IPCC), 1995, p. 109]. Simulations of the tropical tropospheric  $O_3$  distribution are hampered by a lack of knowledge of transport processes such as deep convection, stratosphere-troposphere exchange (STE), and long-range transport from the source regions. The aim of this paper is to study how transport processes affect the tropospheric  $O_3$  distribution over the Indian Ocean.

An important factor in transport studies is the lifetime of  $O_3$ .

<sup>1</sup>Climate Research and Seismology Division/Atmospheric Composition Research, Royal Netherlands Meteorological Institute (KNMI), De Bilt, Netherlands.

<sup>2</sup>Institute for Chemistry of the Polluted Atmosphere (ICG-2), Research Centre Jülich, Jülich, Germany.

<sup>3</sup>Institute for Marine and Atmospheric Research (IMAU), Utrecht University, Utrecht, Netherlands.

<sup>4</sup>National Physical Laboratory, New Delhi, India.

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$O_3$  is photodissociated by short wave solar radiation ( $<340$  nm) into electronically excited O ( $^1D$ ) atoms. Reaction of these with water vapor forms OH radicals. Thus the lifetime of  $O_3$  is basically determined by the amount of water vapor and solar radiation. The lifetime of  $O_3$  increases from 2 to 5 days in the moist tropical marine boundary layer (MBL) to approximately 90 days in the free troposphere [Fishman *et al.*, 1991]. Thus, once  $O_3$  is lifted from the boundary layer, it can be transported far away from its source regions. The tropical continental boundary layer (CBL) is an important source region of photochemically produced  $O_3$ . Biomass burning and fossil fuel combustion generate CO, CH<sub>4</sub>, and nonmethane hydrocarbons. In a NO<sub>x</sub>-rich environment the oxidation of these compounds produces  $O_3$  [Crutzen, 1974; Chameides, 1978; Fishman *et al.*, 1979; Crutzen *et al.*, 1979, 1985; Logan *et al.*, 1981; Greenberg *et al.*, 1984; Koppmann *et al.*, 1997; Chatfield *et al.*, 1998]. Once it is vented from the CBL, this  $O_3$  can be transported over large distances, as is shown by satellite and sounding measurements [Krishnamurti *et al.*, 1993; Fishman *et al.*, 1990, 1991; Piotrowicz *et al.*, 1989; Baldy *et al.*, 1996; De Laat *et al.*, 1999; Thompson *et al.*, 1996; Chatfield *et al.*, 1998; Taupin *et al.*, 1999]. The tropical MBL, on the other hand, is a sink region for  $O_3$ . Kley *et al.* [1996] have measured very low  $O_3$  mixing ratios over the equatorial Pacific in the MBL as well as in the upper troposphere. Lifting of  $O_3$ -poor MBL air, with additional  $O_3$  depletion that may occur in clouds, may have caused these minima [Kley *et al.*, 1996]. Similar minima have been found over the Indian Ocean due to lifting of  $O_3$ -poor MBL air by convective cells of the Intertropical Convergence Zone (ITCZ) [De Laat *et al.*, 1999; Taupin *et al.*, 1999]. An

important source of  $O_3$  in the tropics is also lightning. Since lightning creates a  $NO_x$ -rich environment in the atmosphere,  $O_3$  production can occur [Pickering *et al.*, 1991; Price and Rind, 1994]. Lightning activity is mainly concentrated over the summer hemisphere landmasses in the tropics [Price and Rind, 1994]. Some photochemical ozone production also takes place in the tropical uppermost troposphere [Folkins *et al.*, 1999].

The stratosphere is a third source region through stratosphere-troposphere exchange (STE) by, for example, tropopause folds [Shapiro, 1980]. Exchange can also occur due to filamentation around the subtropical jet (STJ) [Appenzeller *et al.*, 1996; Appenzeller and Holton, 1997]. Shear-induced differential advection can cause tracer surfaces to stretch which leads to quasi-horizontal laminae-like tracer structures [see Appenzeller and Holton, 1997, Figure 1; Ambaum, 1997]. The mechanisms causing a tropopause fold and filamentation are, in principle, the same. The reason that we make a distinction between the two STE exchange mechanisms is because in foldings vertical tilting is important, whereas for the filaments stretching is more important so that they stay close to the tropopause. The mechanism we describe as filamentation is quasi-horizontal, chaotic mixing on a theta surface, which, in principle, conserves both theta and PV. Lastly, vertical shear in the strong shear zone near the STJ, by exceeding the Richardson number, can lead to clear-air turbulence (CAT), which is the third mechanism we distinguish [Pepler *et al.*, 1998; Shapiro, 1978; Kennedy and Shapiro, 1980]. The transport further into the tropics is again differential advection. STE is so far believed to play only a minor role in the tropical tropospheric  $O_3$  budget [Holton and Lelieveld, 1996]. However, stratospheric intrusions have been found near the edges of the tropics, associated with tropopause folds in the vicinity of the subtropical jet stream [Fabian and Pruchniewicz, 1977; Randriambelo *et al.*, 1999; Baray *et al.*, 1998]. Apart from the studies from Pacific Exploratory Mission-West (PEM-West) and Measurements of Ozone by Airbus In-Service Aircraft (MOZAIC) [Newell *et al.*, 1996, 1999; Browell *et al.*, 1996; Wu *et al.*, 1997] and PEM-Tropics [Fenn *et al.*, 1999], other studies also indicated stratospheric influences in the tropical troposphere [Krishnamurti *et al.*, 1993; Kley *et al.*, 1996; Suhre *et al.*, 1997; Cammas *et al.*, 1998; Taupin *et al.*, 1999]. Taupin *et al.* [1999] found stratospheric intrusions into the upper troposphere near the edge of the tropics over the southern Indian Ocean during March and the austral winter (JJA). How far such stratospheric influences extend in the deep tropics needs to be studied in more detail.

To study the influence of transport on tropical tropospheric  $O_3$  profiles, we analyzed  $O_3$  and relative humidity (RH) profiles from the 1998 Indian Ocean Experiment First Field Phase (INDOEX FFP). To analyze air mass origins, we used back trajectories and meteorological data from the European Centre for Medium-Range Weather Forecasts (ECMWF).

## 2. Measurements

During February–March 1998, 14  $O_3$  and radiosondes were launched from the Indian research vessel *Sagar Kanya* over an extensive area of the Indian Ocean. A preliminary overview of the data is given by Mandal *et al.* [1999]. The ship track and location of the soundings are shown in Figure 1. The date and location of each sounding are presented in Table 1.

The sondes used were balloon-borne Electrochemical Concentration Cell (ECC)  $O_3$  sondes (Model 1z, En-Sci Corp., Boulder, Colorado) coupled to Väisälä radiosondes (Model

RS80-15 H, Väisälä, Finland). The uncertainties in the temperature and pressure measurements below 20 km are  $\pm 0.3^\circ\text{C}$  and  $\pm 0.5$  hPa, respectively. The accuracy of the humidity sensor (HUMICAP-H) varies from  $\pm 2\%$  near the surface to  $\pm 15$ – $30\%$  between 5 and 15 km altitude [Kley *et al.*, 1997]. Above 15 km altitude the performance of the sensor is not reliable anymore [Kley *et al.*, 1997]. The accuracy of the  $O_3$  sensor varies from  $\pm 1$ – $2$  ppbv below 5 km to  $\pm 5$  ppbv at 10 km and  $\pm 20$  ppbv at 20 km altitude [Smit *et al.*, 1994, 1998].

We categorized soundings according to their distance from the ITCZ (Table 1). The ITCZ position is estimated from the ECMWF vertical wind analyses at 500 hPa at a  $1^\circ \times 1^\circ$  horizontal resolution. The latitude range given for the ITCZ reflects its width and not the uncertainty in its determination. The categories are (I) far from the ITCZ ( $10^\circ$  or more), (II) just north of the ITCZ (approximately  $2^\circ$ – $5^\circ$ ), (III) just south of the ITCZ (approximately  $2^\circ$ – $5^\circ$ ), and (IV) in or close to the ITCZ (within  $1^\circ$ ).

## 3. Trajectory Model

Trajectories were calculated with the Royal Netherlands Meteorological Institute (KNMI) trajectory model [Scheele *et al.*, 1996]. The model uses ECMWF 6-hour forecast wind fields to calculate the displacement of an air parcel over a 1-hour model time step. Forecast data are used instead of analyzed data in order to rule out imbalances due to data assimilation in the model analyses. The data are available at 31 hybrid sigma-p model levels at a  $1^\circ \times 1^\circ$  horizontal resolution. For more details we refer to the User Guide to ECMWF Products 2.1 (1995, available from ECMWF).

For all calculations we applied clustered 5-day backward kinematic trajectories. We calculated multiple trajectories starting from 8 corners and 19 midpoints of the edges of a three-dimensional (3-D) box around the air mass of interest. The length, width, and height of the box are  $1^\circ$ ,  $1^\circ$ , and 40–80 hPa, respectively. This cluster technique gives an impression of the reliability of the trajectories.

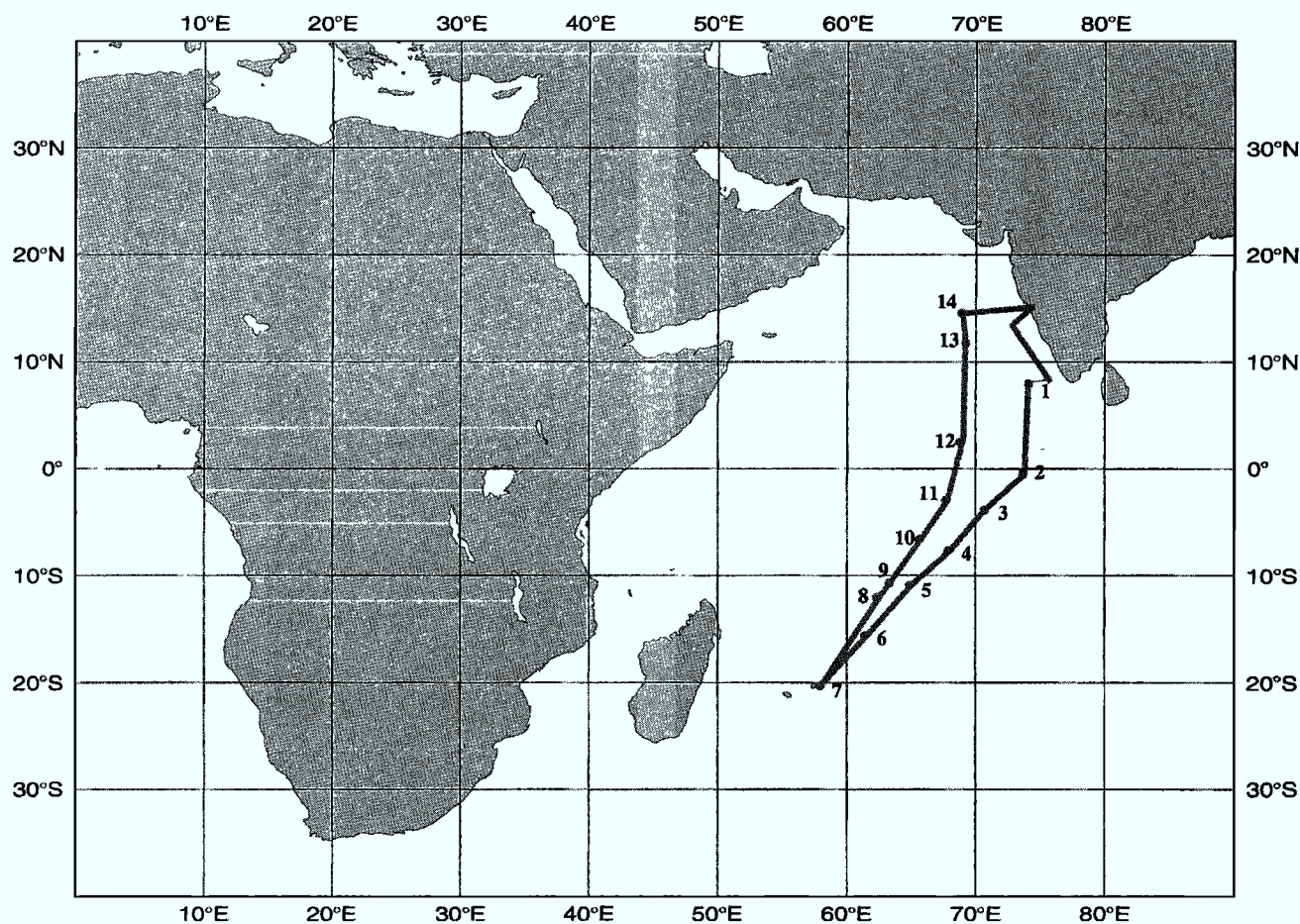
It cannot be excluded that the ECMWF analyses are inaccurate. The tropics is a data-sparse region, especially over the oceans, where the dynamics calculations are difficult due to a lack of a large-scale dynamical balance (such as the geostrophic balance for the extratropics) [Heckley, 1985]. ECMWF data postprocessing, as well as conversion from spherical harmonics space to grid point space, can also introduce errors. We believe, however, that ECMWF provides the best quality and finest resolution analyses around the world. In this study, we used data with a resolution similar to that of the model simulations. At this resolution the model resolves the larger convective cells of the ITCZ, although it probably underestimates the vertical motions herein. Subgrid-scale convection is not captured. Thus the actual path of an air parcel becomes unreliable if the trajectory passes through a convective area. This becomes important close to the ITCZ since individual turrets in the ITCZ are smaller than the model resolution. National Oceanic and Atmospheric Administration (NOAA) IR satellite pictures coincident with the trajectory locations were used to rule out the influence of convection as much as possible.

## 4. Analyses and Discussion of the Sounding Data

### 4.1. Description of the Soundings

Figure 2 shows the categorized profiles. In general, the profiles consist of multiple layers characterized by distinct maxima





**Figure 1.** Shiptrack with the positions of the  $O_3$  soundings. The first and last soundings were launched on February 23 and March 28, 1998, respectively.

and minima. Interesting features are very dry midtropospheric  $O_3$  maxima (profiles 6, 7, and 12), upper tropospheric  $O_3$ -rich laminae (profiles 10–14), and upper tropospheric  $O_3$  minima (e.g., in profiles 6 and 12). A detailed analysis follows in the sections below. In soundings just north of the ITCZ the lower troposphere has a layered structure superimposed on a 20-

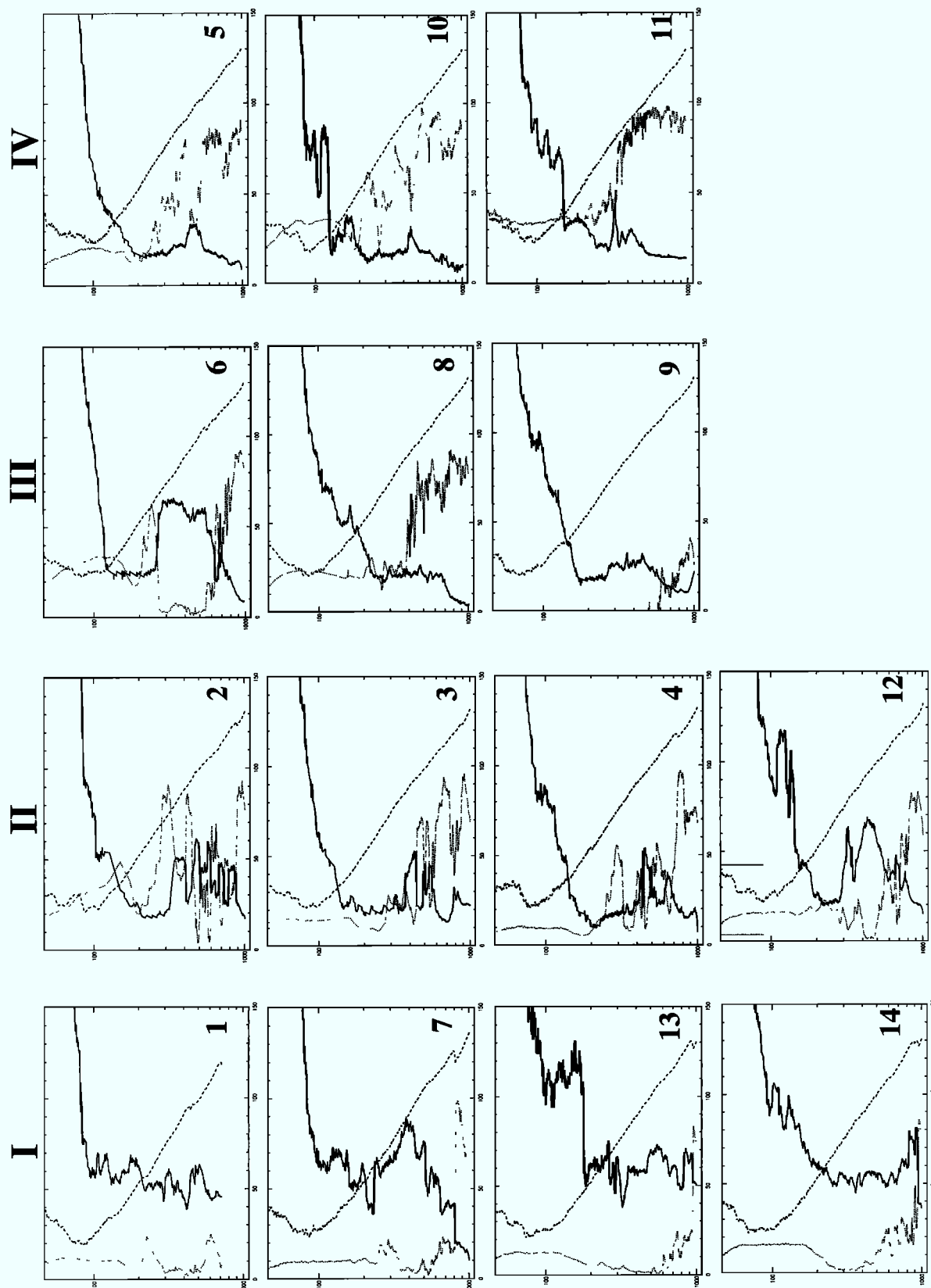
ppbv  $O_3$  background.  $O_3$  and RH are strongly anticorrelated. These layers are similar to those found in the MOZAIC and PEM data sets [Newell *et al.*, 1996, 1999; Browell *et al.*, 1996]. The average altitude of their layers was 5–6 km, to be compared to our altitude range, which was from just above the MBL to 7–8 km. The layer thickness is similar (about 1 km).

The MBL  $O_3$  mixing ratios are generally low (10–20 ppbv), although slightly higher than those observed in 1995 [De Laat *et al.*, 1999]. Values as low as those over the equatorial Pacific [Kley *et al.*, 1996] are not observed here. MBL mixing ratios are enhanced to 40–50 ppbv close to India (soundings 1, 13, and 14). Trajectory analysis (for example, that for profile 13, see Plate 1a) shows that these elevated  $O_3$  mixing ratios are due to photochemical  $O_3$  production from Indian pollution outflow. A second outflow layer (800–900 hPa) overlies the MBL as shown by the trajectories of profile 13 in Plate 1b. The  $O_3$  mixing ratios in this layer are even higher (70 ppbv) due to the longer lifetime of  $O_3$  once it is in the free troposphere. The “background” free tropospheric  $O_3$  mixing ratios are highest in the soundings close to the Indian subcontinent (50–80 ppbv). This is consistent with previous studies, which showed that the free troposphere can be influenced by continental boundary layer (CBL) pollution as a result of deep convective mixing [Thompson *et al.*, 1996; Baldy *et al.*, 1996; Randriambelo *et al.*, 1999; Taupin *et al.*, 1999; De Laat *et al.*, 1999]. The lowest MBL  $O_3$  mixing ratios appear in the soundings just south of the

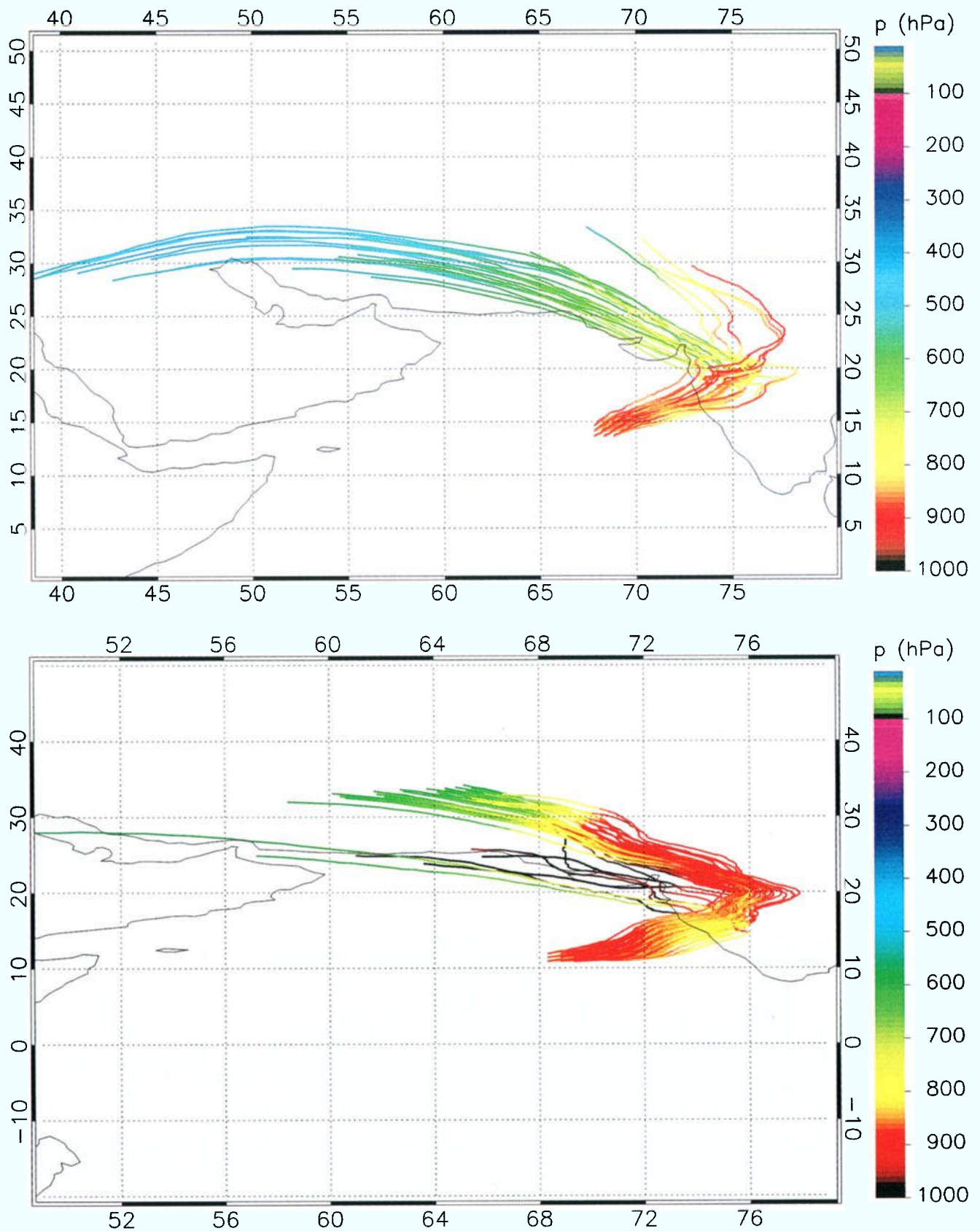
**Table 1.** Locations and Dates of the Soundings

Sounding	Date	Latitude	Longitude	ITCZ	Category
1	Feb. 23, 1998	8.0°N	74.0°E	3°–6°S	I
2	March 2, 1998	0.2°S	73.5°E	2°–5°S	II
3	March 4, 1998	4.3°S	70.2°E	8°–12°S	II
4	March 6, 1998	8.1°S	67.3°E	12°–17°S	II
5	March 8, 1998	11.1°S	64.5°E	12°–15°S	IV
6	March 10, 1998	16.1°S	61.0°E	12°–14°S	III
7	March 13, 1998	20.1°S	57.3°E	6°–8°S	I
8	March 18, 1998	12.4°S	62.5°E	4°–8°S	III
9	March 19, 1998	11.0°S	63.3°E	2°–7°S	III
10	March 20, 1998	7.1°S	65.2°E	3°–10°S	IV
11	March 22, 1998	3.0°S	67.3°E	5°S–2°N	IV
12	March 24, 1998	2.4°N	68.5°E	2°–10°S	II
13	March 27, 1998	11.4°N	68.8°E	0°–5°S	I
14	March 28, 1998	14.1°N	68.3°E	0°N–9°S	I

The soundings were launched either at the beginning of the afternoon or in the evening (local time). Also indicated is the position of the Intertropical Convergence Zone (ITCZ) and the categories into which the soundings were partitioned (see text section 2).

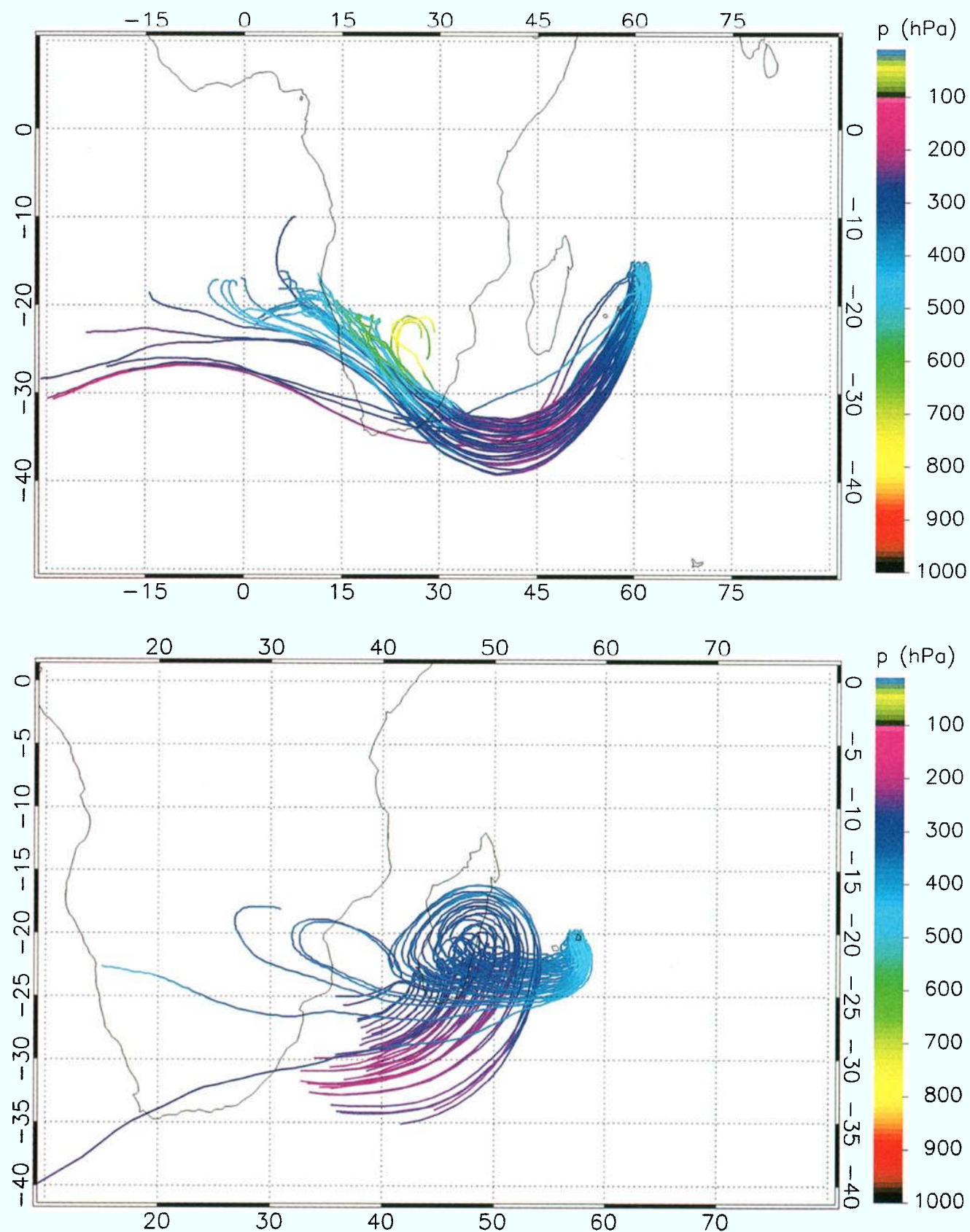


**Figure 2.** O<sub>3</sub>, RH, and temperature profiles. Solid lines: O<sub>3</sub> (in ppbv), dotted lines: RH (in %), dashed lines: temperature (in K) offset by 170 K in order to fit into the plot. The horizontal axis runs from 0 to 150 (units: ppbv, %, and K). The vertical axis denotes the pressure from 1010 to 50 hPa. The soundings are classified into four categories depending on their position relative to the ITCZ: (I), from the ITCZ; (II), just north of the ITCZ; (III), in the ITCZ; (IV), just south of the ITCZ. The tropopause is indicated by the temperature minimum.



**Plate 1.** Five-day back trajectories from a  $1^\circ \times 1^\circ$  box surrounding the position of sounding 13. Trajectories end on March 27 at the following levels: (a) 920, 940, 960 hPa (in the MBL) and (b) 800, 820, 840, 860, 880 hPa (layer just above the MBL).





**Plate 2.** Five-day back trajectories from a  $1^\circ \times 1^\circ$  box surrounding the ozone maxima of soundings 6, 7, and 12. Trajectories end on the following days and levels: (a) March 10 (sounding 6) at 360, 380, 400, 420, 440, 460, 480 hPa. (b) March 13 (sounding 7) at 380, 400, 420, 440, 460 hPa. (c) March 24 (sounding 12) at 410, 430, 450, 470, 490 hPa.

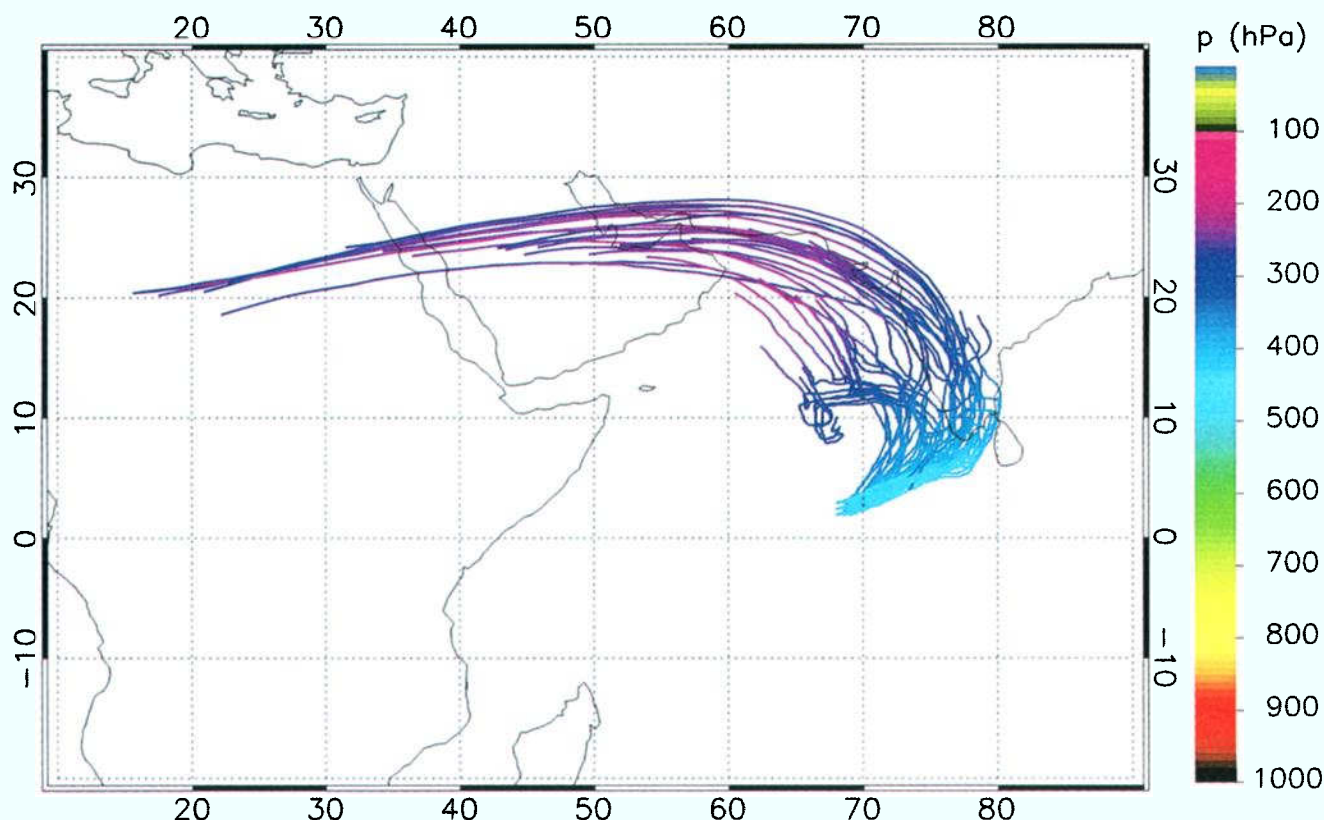


Plate 2. (continued)

ITCZ and in the ITCZ itself, reaching near-zero levels in soundings 4 and 8. We now analyze some of the general features in more detail.

#### 4.2. Dry, Midtropospheric O<sub>3</sub> Maxima

The midtropospheric maxima in the profiles close to or in the ITCZ (5, 10, and 11) are a result of the convective activity. The lifetime of O<sub>3</sub> is short in the MBL. Convection vents O<sub>3</sub>-poor air from the MBL to the upper troposphere [Lelieveld and Crutzen, 1994; Roelofs *et al.*, 1997; Smit *et al.*, 1991]. This O<sub>3</sub>-poor air at upper levels is only slowly mixed back to lower levels by subsidence in the cloud-free part of the ITCZ. At midtropospheric levels the convective detrainment is often small. Thus O<sub>3</sub>-poor air can develop in the upper troposphere and in the MBL, while the middle troposphere remains relatively unaffected. This results in a relative O<sub>3</sub> maximum in the middle troposphere.

The maxima in soundings 6, 7, and 12, which were released outside the ITCZ, have a different origin. These O<sub>3</sub> maxima are superimposed on background values of 20 ppbv and are extremely dry (RH < 5%). This dryness is also found in the ECMWF specific humidity analyses, of which a representative example is shown in Figure 3 for profile 6. The dry layer is part of a synoptic structure that slopes upward toward the STJ. The specific humidity analyses for profiles 7 and 12 are similar, with a dry layer at the location of the midtropospheric maximum. The fact that the model simulates a layer similar to that measured lends extra confidence to the back trajectory analyses, which are based on wind fields from this model.

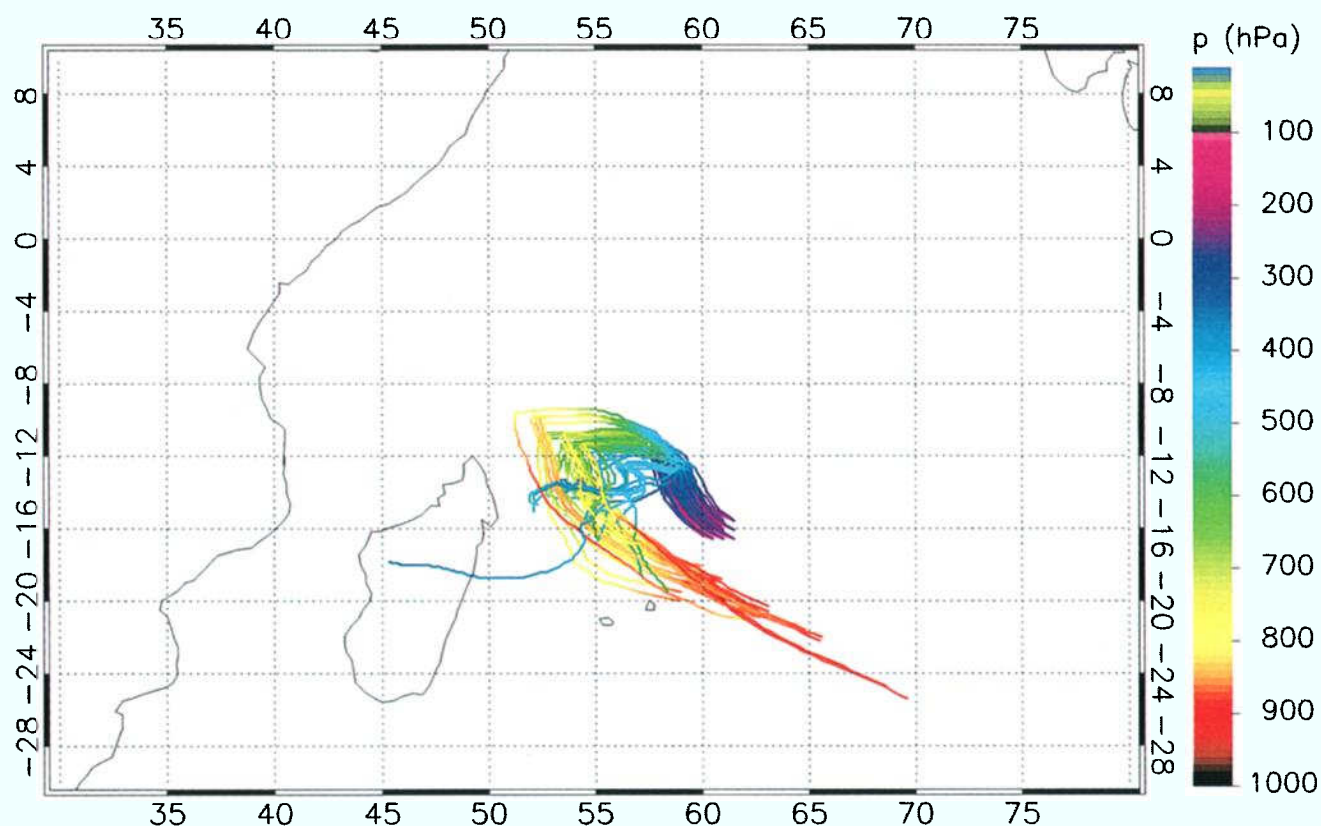
The combination of high O<sub>3</sub> and low RH might be due to either a stratospheric origin or to convectively lifted CBL air

that has subsided. Back trajectories show that the air masses in the maxima in soundings 6 and 7 (Plates 2a–2c) originated from the upper troposphere at 30°–40°S. Those for sounding 12 originated from 20°–30°N. These are exactly the locations of the STJ according to the horizontal wind analyses at 200 hPa (for example, Figure 4). A large vertical and horizontal shear zone surrounds this STJ, making this a favored region for STE. We therefore conclude that the maxima are most likely due to STE near the STJ. The influence of convective lifting of CBL air over India in sonde 12 was ruled out by comparison with NOAA IR satellite pictures. These showed no convective events during the period that the trajectories passed over India. India was still in the dry monsoon during that period, and large-scale subsidence prevailed. The extreme dryness of the layer in sonde 6 makes it highly unlikely that this air mass comes from the polluted CBL.

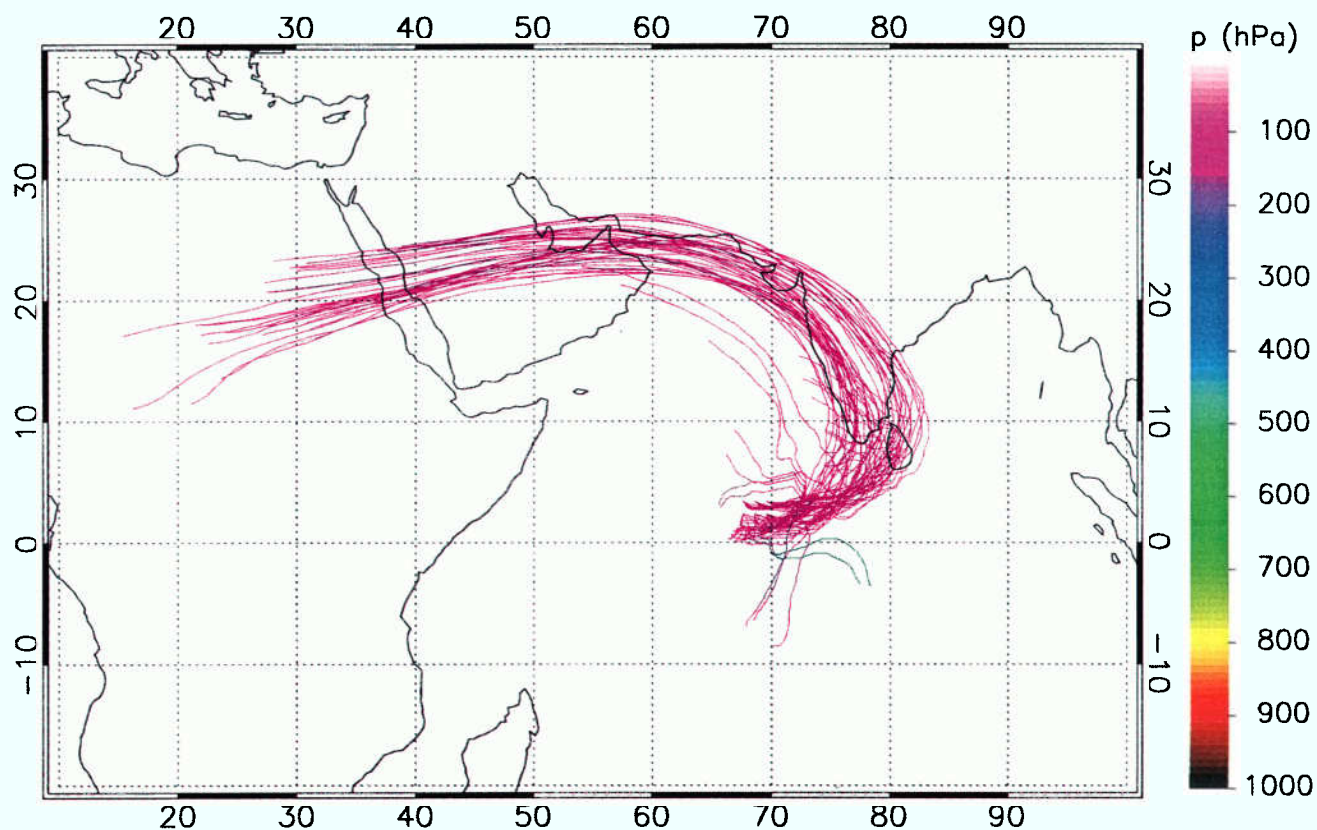
As described in the introduction, the STE can be induced by tropopause folds, shear-induced filament stripping from the STJ, or CAT followed by shear-induced differential advection. We did not find any evidence in the ECMWF data of tropopause folds in the sense of a deep vertically inclined frontal layer with high O<sub>3</sub> and PV and low RH. The explanation for the O<sub>3</sub> maxima is therefore probably one of the latter two STE mechanisms.

The RH decreases as the cluster further descends anticyclonically toward the deep tropics. During this descent the air has most likely mixed with tropospheric air since the O<sub>3</sub> values in the maximum are not purely stratospheric (70–90 ppbv compared to lower stratospheric values typically above 100 ppbv). A stratospheric PV signature along the trajectory could not be found for these O<sub>3</sub> maxima. In profiles 6 and 12, O<sub>3</sub>



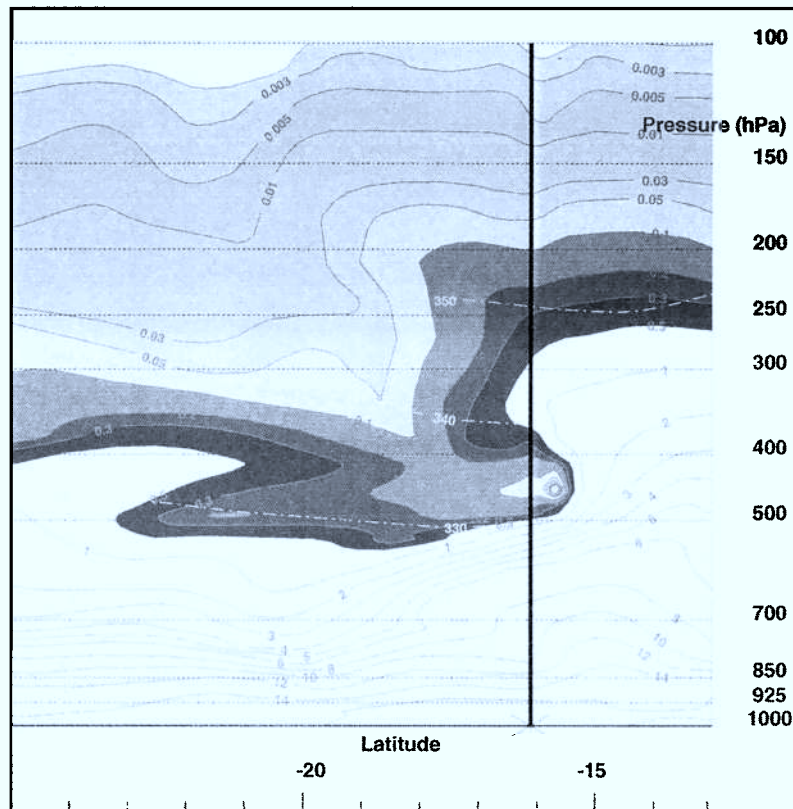


**Plate 3.** Five-day back trajectories from a  $1^\circ \times 1^\circ$  box surrounding the upper tropospheric ozone minimum of sounding 6. Trajectories end on March 10 at the following levels: 160, 180, 200, 220, 240 hPa.



**Plate 4.** Backward trajectories starting from the upper tropospheric laminae of sounding 12. Trajectories from March 19–24 from 100 to 160 hPa at 10 hPa intervals.





**Figure 3.** ECMWF specific humidity cross section (in g/kg) for profile 6 on March 10, 0600 UTC at 61°E. The position of the profile is indicated by the vertical line.

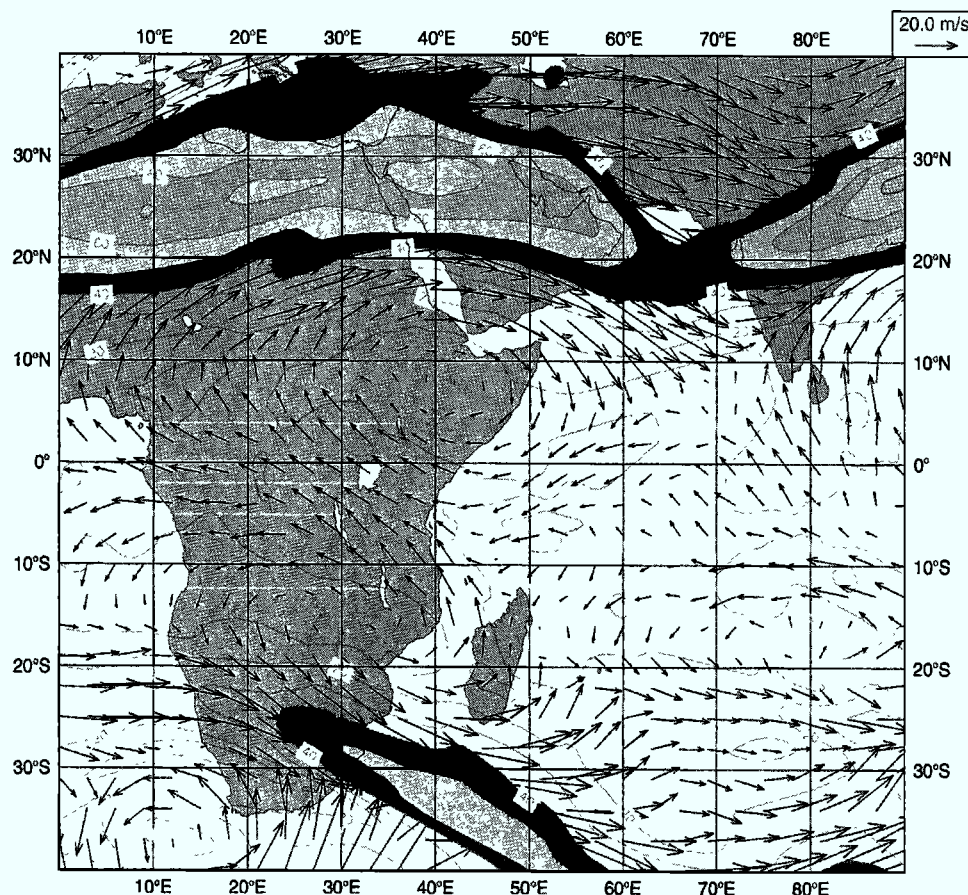
minima, with mixing ratios similar to near-surface values, overlie the midtropospheric maxima. Back trajectories from profile 6 (Plate 3) indicate that these air masses were lifted from the  $O_3$ -poor MBL. The ECMWF horizontal wind analysis (not shown) indicates that this vertical transport is caused by a cyclone-like perturbation, which involves rapid upward transport due to latent heat release. *De Laat et al.* [1999] show that these kind of upward motions are resolved (but probably underestimated) by the ECMWF data.

#### 4.3. Upper Tropospheric Laminae

Trajectory analyses (of which a representative example is shown for profile 12 in Plate 4) indicate that most laminae originate near the STJ. After spending some time near the jet stream, the trajectories curve anticyclonically toward the equator. Similar to the midtropospheric maxima, we therefore expect that the  $O_3$ -rich laminae are the result of STE around the STJ. The analysis of the potential temperature ( $\theta$ ) along one of the trajectories of Plate 4 (at 110 hPa) corroborates the hypothesis of a stratospheric origin of these laminae (Figure 5). Since the tropical tropopause more or less coincides with the 380 K surface, the analysis shows that  $\theta$  overshoots the tropopause value during the first few days. However, it must be kept in mind that the ECMWF vertical velocity field may have errors, which influence the  $\theta$  evolution along the trajectory. Although the absolute change in  $\theta$  ( $>10$  K in  $<10$  hours) is rather large, we have some confidence in the general trend of  $\theta$  to decline over several days. Part of the temporal variations in  $\theta$  is due to vertical interpolation errors of  $\theta$  to the trajectory position.

Another indication of a stratospheric origin of the laminae of profile 12 is found in the PV map at 100 hPa (Figure 6a). A patch of typically stratospheric PV values is shown near the measurement site (the tropopause is defined by the 2 PVU contour). Laminae with stratospheric PV (and  $O_3$ ) values appear to be stripped off the STJ (indicated also by following the filament a few days back, not shown) and transported anticyclonically along the trajectory paths over Sri Lanka and southern India to the measurement location. This stripping followed by transport toward the deep tropics is also observed in PV maps pertaining to other laminae. An example is sounding 7. Although not as pronounced as in the other soundings, a layer with elevated  $O_3$  mixing ratios (70 ppbv) appears between 100 and 150 hPa. The corresponding PV map at 100 hPa (Figure 6b) shows a strongly meandering STJ. This meandering is accompanied by laminae stripping, evidenced by several tongues of stratospheric PV values (1.5–2 PVU) penetrating into the deep tropics and to the measurement site. Comparing the  $O_3$  mixing ratios in laminae in the deep tropics to those close to the STJ, it seems that during transport mixing was not as important as for the midtropospheric ozone maxima. The layers can probably retain their identity for a long time because photochemical  $O_3$  destruction is small in the dry upper troposphere and there may even be some  $O_3$  production. We do not believe, however, that local production plays a major role in causing these layers in view of the evidence for STE presented above.

Another explanation for  $O_3$ -rich laminae might have been transport from the polluted CBL. However, it is most unlikely that low-level pollution creates stable layers with  $O_3$  mixing



**Figure 4.** ECMWF horizontal wind and wind speed on March 10 (sounding 6) at 200 hPa. The Northern Hemisphere subtropical jet stream is stronger than the Southern Hemisphere jet stream.

ratios up to 120 ppbv in the upper troposphere. None of the trajectories showed any indication of recent contact with the polluted CBL since the entire cluster originates from the region of the STJ. The cluster remains in the upper troposphere during the 5-day period. Moreover, NOAA IR satellite pictures showed no convective events during the period that the trajectories passed over source regions such as India. This points to a stratospheric rather than a continental origin of the laminae.

Only a few of the trajectories from the laminae do not originate near the STJ (not shown). Instead, these seem to originate from the ITCZ. Yet also for these layers there are indications in the PV maps (not shown) that the air masses originate, at least partly, from the stratosphere.

## 5. Conclusions and Discussion

The picture that emerges from our analysis is that  $O_3$  profiles over the Indian Ocean are often composed of several distinct air masses that cause alternating maxima, minima, and laminae in ozone. The observed profiles display the following features:

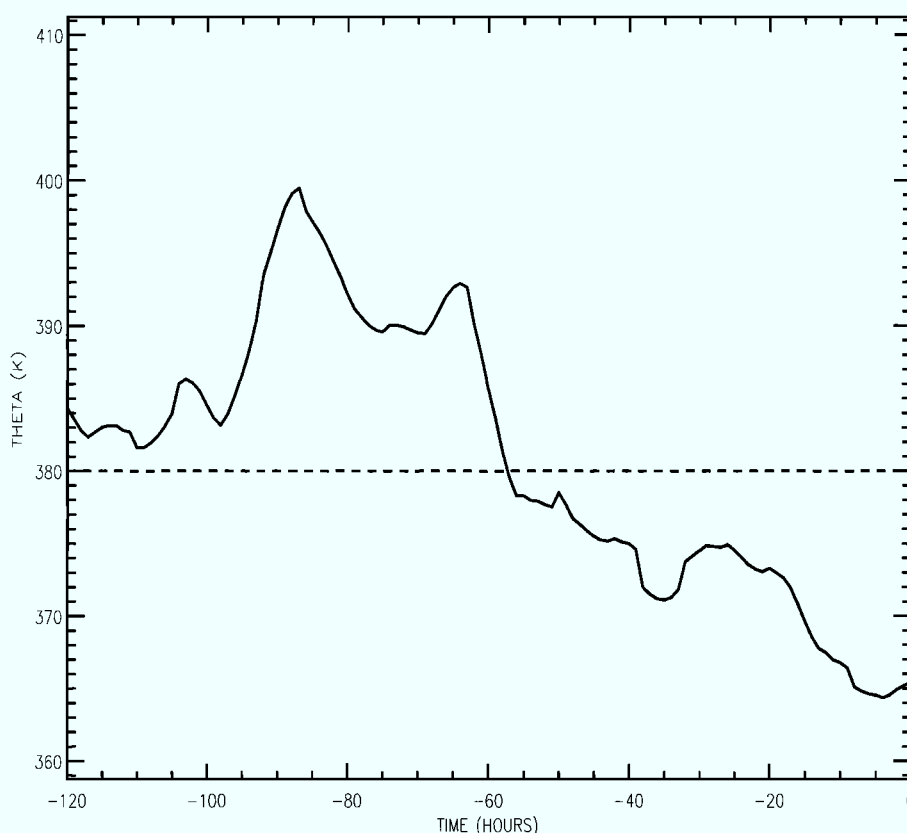
1. MBL mixing ratios are enhanced to 40–50 ppbv close to India due to outflow of pollution from the continent. MBL mixing ratios away from the Asian continent are relatively low (10–20 ppbv) consistent with efficient photochemical  $O_3$  destruction. Convective lifting from the MBL can cause upper tropospheric  $O_3$  minima.

2. There are elevated  $O_3$  mixing ratios (50–80 ppbv) above the MBL originating from low-level outflow of polluted air from the Asian continent.

3. There are very dry midtropospheric maxima (60–90 ppbv) in the middle troposphere. We have presented evidence that this air most likely has a stratospheric origin.

4. There are high to very high  $O_3$  mixing ratios (up to 120 ppbv) in laminae just below the tropopause. Also for these layers we have presented evidence of a stratospheric origin, with the STJ playing a central role in the exchange mechanism.

Stratospheric inputs have also been found in the tropics by Newell *et al.* [1996, 1999], Browell *et al.* [1996], Wu *et al.* [1997], and Fenn *et al.* [1999], but the repeated presence of stratospheric laminae just below the tropopause is a remarkable new finding of this study. Similar laminae are present in Central Equatorial Pacific Experiment (CEPEX) profiles over the central Pacific [Kley *et al.*, 1996], but they were not discussed by these authors. The laminae appear in more than a third of the profiles. We have shown with back trajectory analyses that the laminae originate from the vicinity of the STJ. We hypothesize that the STJ plays a central role in this STE mechanism either by directly shear-induced differential advection, which causes filamentation, or by CAT followed by differential advection. In principle, tropopause folds could also play a role, but we did not find evidence for these in the ECMWF data. Although the ECMWF data have a limited vertical and horizontal resolution, observed tropopause folds are often well captured by the



**Figure 5.** Potential temperature evolution along a trajectory from the cluster of sounding 12 starting in a layer at 110 hPa (see Plate 4). Time equal to 0 is the starting point of the trajectory (March 24), and time equal to  $-120$  is the endpoint of the trajectory (March 19).

ECMWF model. We therefore believe it is more likely that STE has taken place due to shear-induced differential advection possibly preceded by CAT. ECMWF PV maps illustrate the stripping of laminae from the STJ in several cases.

However, stratospheric PV signatures along the trajectories were not always found. This does not necessarily imply that the air cannot be of stratospheric origin. There can be other reasons for a lack of conservation of a PV signal along the trajectory. If stratospheric air is mixed into the troposphere by CAT, then a pronounced PV signal would be lacking. Furthermore, even if the trajectory is perfect, one should realize that the PV is a grid box value. Thus, even if high PV laminae are present in the grid box, the PV signal will probably be suppressed by grid box averaging. Another reason could be that during transport the PV signature disappears rapidly due to absorption of solar radiation. In this case the chemical signature would remain unaffected. This may be especially important for thin layers in the tropics.

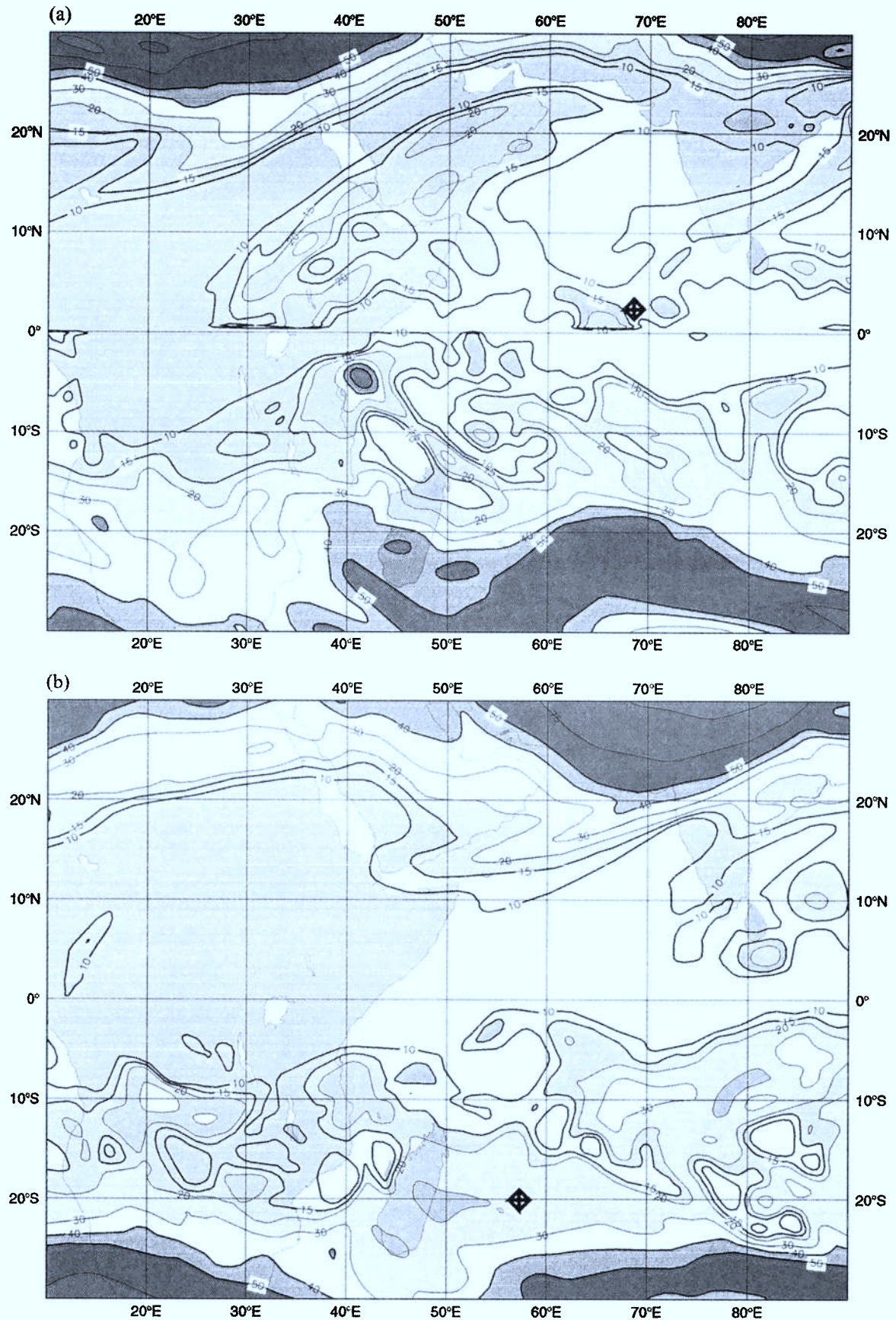
Only a few of the laminae do not seem to originate from the STJ but rather from the ITCZ. Yet also for these layers there are indications of a stratospheric origin in the PV maps (not shown). This implies that either the trajectories are wrong, which is likely near convection, or the exchange mechanism involves the ITCZ rather than the STJ. Overshooting clouds in the ITCZ can entrain stratospheric air into the turret [Danielsen, 1982]. Radiative and evaporative cooling cause the

mixed air mass to sink back to the troposphere. Overshooting clouds also generate buoyancy waves in regions around the ITCZ [Potter and Holton, 1995; Danielsen, 1993]. These can cause STE, although this does not explain the observed horizontal orientation of the laminae. Finally, ozone production due to lightning associated with deep convection may also play a role. Further analyses will be needed in future.

Our results do not point to preferred latitudes for these stratospherically influenced layers. Rather, we have shown that the upper tropospheric laminae can be transported quasi-horizontally quite far away from the STJ into the deep tropics. To our knowledge, stratospheric influences in the form of thin layers with high  $O_3$  concentrations close to the tropical tropopause have not been discussed in the literature before. The  $O_3$  mixing ratios in the laminae close to the equator were as high as those close to the STJ, suggesting that mixing is not as important as for the midtropospheric  $O_3$  maxima, in which mixing ratios were lower than those of the upper tropospheric laminae. Probably, the midtropospheric maxima undergo stronger influences from convection and turbulence.

Future work should concentrate on the frequency and distribution of these intrusions into the tropical troposphere and on aspects such as seasonal variability and interannual variability. Statistical analysis and chemistry-transport simulations should be performed in order to estimate the magnitude of the contribution of these intrusions to the total tropospheric  $O_3$  budget in the tropics.





**Figure 6.** ECMWF potential vorticity (in 0.1 PVU, 1 PVU =  $10^{-6} \text{ m}^2 \text{ K kg}^{-1} \text{ s}^{-1}$ ) at 100 hPa. The location of the profiles is given by a symbol. (a) March 24, 0600 UTC (sounding 12; 2.4°N/68.5°E) and (b) March 13, 1200 UTC (sounding 7; 20.1°S/57.3°E).

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H. Kelder, P. F. J. van Velthoven, and M. Zachariasse, Climate Research and Seismology Division/Atmospheric Composition Research, KNMI, P.O. Box 201, 3730 AE De Bilt, Netherlands. (zacharia@knmi.nl)

J. Lelieveld, Institute for Marine and Atmospheric Research (IMAU), Utrecht University, Princetonplein 5, 3584 CC Utrecht, Netherlands.

T. K. Mandal, National Physical Laboratory, Dr. K. S. Krishnan Road, New Delhi 110-012, India.

H. G. J. Smit, Institute for Chemistry of the Polluted Atmosphere (ICG-2), Research Centre Jülich, P.O. Box 1913, D-52425 Jülich, Germany.

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