

# Error estimation for soil moisture measurements with cosmic-ray neutron sensing and implications for rover surveys

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## Motivation

- The aboveground epithermal neutron intensity is effectively determined by surrounding hydrogen, enabling field-scale soil moisture measurements
- The uncertainty of soil moisture measurements with cosmic-ray neutron sensing (CRNS) among other sources depends on the uncertainty in poisson distributed neutron counts: decreasing soil moisture corresponds to decreasing neutron intensity and increasing uncertainty in neutron counts  
→ Reduced uncertainty with more effective detectors, arrays of detectors (e.g., CRNS roving) or aggregation over long time windows (e.g., 12 or 24 h)
- We present an easy-to-apply method for assessing the soil moisture uncertainty from neutron counts, compare it to a computationally intensive Monte Carlo approach and elaborate on implications for the planning and evaluation of CRNS rover surveys

## Uncertainty from neutrons

Cosmic-ray neutrons can be converted to soil moisture via:

$$\theta_v = \rho_{bd} \left[ a_0 \left( \frac{N_{cor}}{N_0} - a_1 \right)^{-1} - a_2 - \theta_{off} \right]$$

Desilets et al., (2010)

$\theta_v$  = volumetric soil moisture [m<sup>3</sup>/m<sup>3</sup>]

$\rho_{bd}$  = soil bulk density [g/cm<sup>3</sup>]

$N_{cor}$  = pressure, humidity and incoming flux corrected neutron counts [cts]

$N_0$  = calibration parameter, usually calibrated with reference soil moisture [cts]

$\theta_{off}$  = constant water pools (e.g., soil organic carbon, lattice water) [g/g]

$a_i$  = fitting parameters obtained by Desilets et al. (2010)

The uncertainty in neutron counts is defined by  $\sqrt{N}$  and for error estimation must be propagated to corrected neutron counts:

$$\sigma_N = s\sqrt{N}$$

$s$  = product of corrections for pressure, humidity and incoming flux

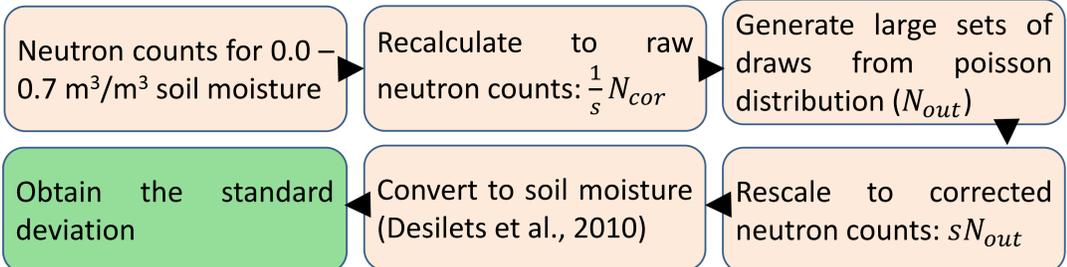
$\sigma_N$  = uncertainty in corrected neutron counts

We used an **analytical Taylor expansion approach** up to 3<sup>rd</sup> polynomial order, considering 6 central moments in the uncertainty distribution (Mekid and Vaja, 2008). Because the neutron count detection statistics converges to a symmetric Gaussian normal distribution, only the 2<sup>nd</sup>, 4<sup>th</sup> and 6<sup>th</sup> moments are relevant:

$$\sigma_{\theta_v} = \rho_{bd} \frac{a_0 N_0}{(N - a_1 N_0)^4} \sqrt{(N - a_1 N_0)^4 + 8\sigma_N^2 (N - a_1 N_0)^2 + 15\sigma_N^4}$$

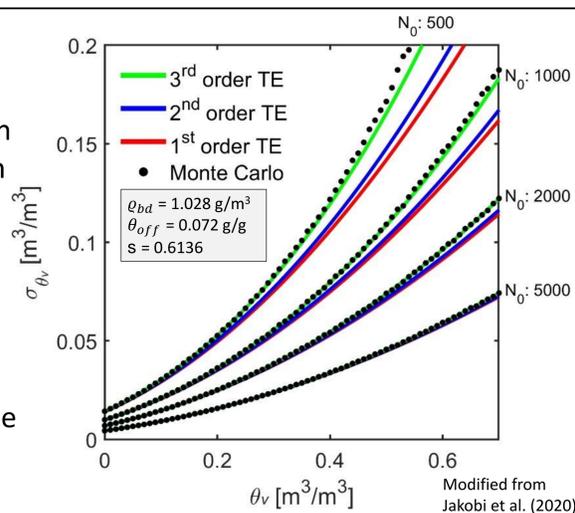
$\sigma_{\theta_v}$  = volumetric soil moisture uncertainty from neutron counts [m<sup>3</sup>/m<sup>3</sup>]

For evaluation we used a **Monte Carlo approach**:

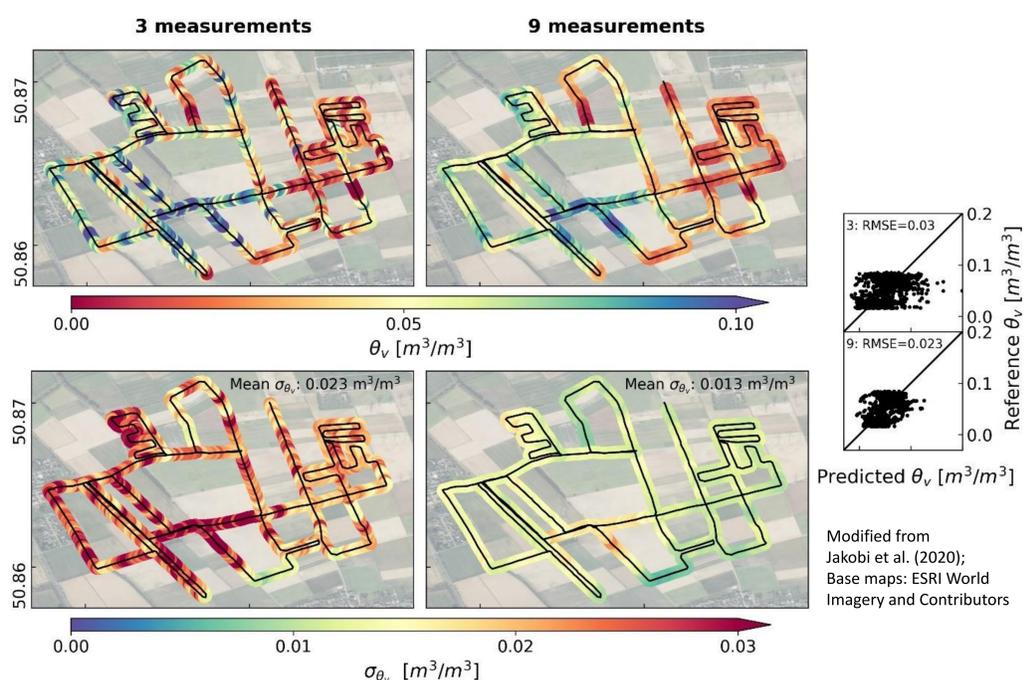


## Method comparison and application

- As expected increasing soil moisture leads to increasing uncertainty in soil moisture estimation
- Aggregation (increasing  $N_0$ ) reduces the uncertainty
- Below  $N_0 = 1000$  slight deviations with high water content due to the asymptote at  $a_1 N_0$

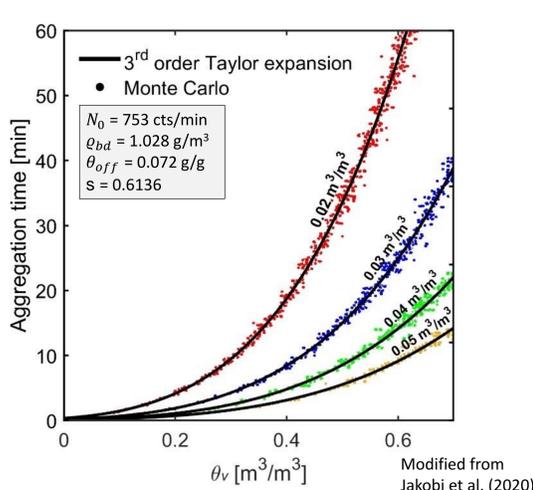


Example of the influence of **aggregation** on the measurement uncertainty



- Clear accuracy improvement due to aggregation  
→ Decreased RMSE and decreased soil moisture uncertainty from neutrons
- Additional sources of uncertainty are not considered here and explain the remaining RMSE:  
→ e.g., uncertainties in soil bulk density, vegetation, roads, inconsistencies between in-situ and CRNS measurements

- With the Jülich CRNS rover and soil moistures below 0.4 m<sup>3</sup>/m<sup>3</sup> an aggregation time up to 10 min is necessary to achieve an uncertainty below 0.03 m<sup>3</sup>/m<sup>3</sup>
- With lower uncertainty requirements the necessary aggregation time can be drastically reduced



## Conclusions

- Measurement uncertainty from neutron counts can be easily estimated with the proposed approach
- With appropriate aggregation the uncertainty can be reduced significantly
- The aggregation length for a roving experiment needs to be carefully selected based on:  
→ rover capabilities; site characteristics; accuracy requirements of the user

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